

Remarks on Conformal Mass and Quantum Mass

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One shows how in certain model situations conformal general relativity corresponds to a Bohmian-Dirac-Weyl theory with conformal mass and Bohmian quantum mass identified.

The article [12] was designed to show relations between conformal general relativity (CGR) and Dirac-Weyl (DW) theory with identification of conformal mass \hat{m} and quantum mass \mathfrak{M} following [7, 9, 11, 25] and precision was added via [21]. However the exposition became immersed in technicalities and details and we simplify matters here. Explicitly we enhance the treatment of [7] by relating \mathfrak{M} to an improved formula for the quantum potential based on [21] and we provide a specific Bohmian-Dirac-Weyl theory wherein the identification of CGR and DW is realized. Much has been written about these matters and we mention here only [1–7, 9–20, 23–28] and references therein. One has an Einstein form for GR of the form

$$S_{GR} = \int d^4x \sqrt{-g} (R - \alpha |\nabla\psi|^2 + 16\pi L_M) \quad (1.1)$$

(cf. [7, 22]) whose conformal form (conformal GR) is an integrable Weyl geometry based on

$$\begin{aligned} \hat{S}_{GR} &= \int d^4x \sqrt{-\hat{g}} e^{-\psi} \times \\ &\times \left[\hat{R} - \left(\alpha - \frac{3}{2} \right) |\hat{\nabla}\psi|^2 + 16\pi e^{-\psi} L_M \right] = \\ &= \int d^4x \sqrt{-\hat{g}} \left[\hat{\phi} \hat{R} - \left(\alpha - \frac{3}{2} \right) \frac{|\hat{\nabla}\hat{\phi}|^2}{\hat{\phi}} + 16\pi \hat{\phi}^2 L_M \right] \end{aligned} \quad (1.2)$$

where $\Omega^2 = \exp(-\psi) = \hat{\phi}$ with $\hat{g}_{ab} = \Omega^2 g_{ab}$ and $\hat{\phi} = \exp(\psi) = \hat{\phi}^{-1}$ (note $(\hat{\nabla}\psi)^2 = (\hat{\nabla}\hat{\phi})^2 / (\hat{\phi})^2$). One sees also that (1.2) is the same as the Brans-Dicke (BD) action when $L_M = 0$, namely (using \hat{g} as the basic metric)

$$S_{BD} = \int d^4x \sqrt{-\hat{g}} \left[\hat{\phi} \hat{R} - \frac{\omega}{\hat{\phi}} |\hat{\nabla}\hat{\phi}|^2 + 16\pi L_M \right]; \quad (1.3)$$

which corresponds to (1.2) provided $\omega = \alpha - \frac{3}{2}$ and $L_M = 0$. For (1.2) we have a Weyl gauge vector $w_a \sim \partial_a \psi = \partial_a \hat{\phi} / \hat{\phi}$ and a conformal mass $\hat{m} = \hat{\phi}^{-1/2} m$ with $\Omega^2 = \hat{\phi}^{-1}$ as the conformal factor above. Now in (1.2) we identify \hat{m} with the quantum mass \mathfrak{M} of [25] where for certain model situations $\mathfrak{M} \sim \beta$ is a Dirac field in a Bohmian-Dirac-Weyl theory as in (1.8) below with quantum potential Q determined via $\mathfrak{M}^2 = m^2 \exp(Q)$ (cf. [10, 11, 21, 25] and note that $m^2 \propto T$ where $8\pi T^{ab} = (1/\sqrt{-g})(\delta\sqrt{-g} \mathfrak{L}_M / \delta g_{ab})$). Then $\hat{\phi}^{-1} = \hat{m}^2 / m^2 = \mathfrak{M}^2 / m^2 \sim \Omega^2$ for Ω^2 the standard conformal

factor of [25]. Further one can write **(1A)** $\sqrt{-\hat{g}} \hat{\phi} \hat{R} = \hat{\phi}^{-1} \sqrt{-\hat{g}} \hat{\phi}^2 \hat{R} = \hat{\phi}^{-1} \sqrt{-g} \hat{R} = (\beta^2 / m^2) \sqrt{-g} \hat{R}$. Recall here from [11] that for $g_{ab} = \hat{\phi} \hat{g}_{ab}$ one has $\sqrt{-g} = \hat{\phi}^2 \sqrt{-\hat{g}}$ and for the Weyl-Dirac geometry we give a brief survey following [11, 17]:

1. Weyl gauge transformations: $g_{ab} \rightarrow \tilde{g}_{ab} = e^{2\lambda} g_{ab}$; $g^{ab} \rightarrow \tilde{g}^{ab} = e^{-2\lambda} g^{ab}$ — weight e.g. $\Pi(g^{ab}) = -2$. β is a Dirac field of weight -1. Note $\Pi(\sqrt{-g}) = 4$;
2. Γ_{ab}^c is Riemannian connection; Weyl connection is $\hat{\Gamma}_{ab}^c$ and $\hat{\Gamma}_{ab}^c = \Gamma_{ab}^c = g_{ab} w^c - \delta_b^c w_a - \delta_a^c w_b$;
3. $\nabla_a B_b = \partial_a B_b - B_c \Gamma_{ab}^c$; $\nabla_a B^b = \partial_a B^b + B^c \Gamma_{ca}^b$;
4. $\hat{\nabla}_a B_b = \partial_a B_b - B_c \hat{\Gamma}_{ab}^c$; $\hat{\nabla}_a B^b = \partial_a B^b + B^c \hat{\Gamma}_{ca}^b$;
5. $\hat{\nabla}_\lambda g^{ab} = -2g^{ab} w_\lambda$; $\hat{\nabla}_\lambda g_{ab} = 2g_{ab} w_\lambda$ and for $\Omega^2 = \exp(-\psi)$ the requirement $\nabla_c g_{ab} = 0$ is transformed into $\hat{\nabla}_c \hat{g}_{ab} = \partial_c \psi \hat{g}_{ab}$ showing that $w_c = -\partial_c \psi$ (cf. [7]) leading to $w_\mu = \hat{\phi}_\mu / \hat{\phi}$ and hence via $\beta = m \hat{\phi}^{-1/2}$ one has $w_c = 2\beta_c / \beta$ with $\hat{\phi}_c / \hat{\phi} = -2\beta_c / \beta$ and $w^a = -2\beta^a / \beta$.

Consequently, via $\beta^2 \hat{R} = \beta^2 R - 6\beta^2 \nabla_\lambda w^\lambda + 6\beta^2 w^\lambda w_\lambda$ (cf. [11, 12, 16, 17]), one observes that $-\beta^2 \nabla_\lambda w^\lambda = -\nabla_\lambda (\beta^2 w^\lambda) + 2\beta \partial_\lambda \beta w^\lambda$, and the divergence term will vanish upon integration, so the first integral in (1.2) becomes

$$I_1 = \int d^4x \sqrt{-g} \left[\frac{\beta^2}{m^2} R + 12\beta \partial_\lambda \beta w^\lambda + 6\beta^2 w^\lambda w_\lambda \right]. \quad (1.4)$$

Setting now $\alpha - \frac{3}{2} = \gamma$ the second integral in (1.2) is

$$\begin{aligned} I_2 &= -\gamma \int d^4x \sqrt{-\hat{g}} \hat{\phi} \frac{|\hat{\nabla}\hat{\phi}|^2}{|\hat{\phi}|^2} = \\ &= -4\gamma \int d^4x \sqrt{-\hat{g}} \hat{\phi}^{-1} \hat{\phi}^2 \frac{|\hat{\nabla}\beta|^2}{\beta^2} = \\ &= -\frac{4\gamma}{m^2} \int d^4x \sqrt{-g} |\hat{\nabla}\beta|^2, \end{aligned} \quad (1.5)$$

while the third integral in the formula (1.2) becomes **(1B)** $16\pi \int \sqrt{-g} d^4x L_M$. Combining now (1.4), (1.5), and **(1B)** gives then

$$\begin{aligned} \hat{S}_{GR} &= \frac{1}{m^2} \int d^4x \sqrt{-g} \left[\beta^2 R + 6\beta^2 w^\alpha w_\alpha + \right. \\ &\quad \left. + 12\beta \partial_\alpha \beta w^\alpha - 4\gamma |\hat{\nabla}\beta|^2 + 16\pi m^2 L_M \right]. \end{aligned} \quad (1.6)$$

We will think of $\hat{\nabla}\beta$ in the form **(1C)** $\hat{\nabla}_\mu\beta = \partial_\mu\beta - w_\mu\beta = -\partial_\mu\beta$. Putting then $|\hat{\nabla}\beta|^2 = |\partial\beta|^2$ (1.6) becomes (recall $\gamma = \alpha - \frac{3}{2}$)

$$\hat{S}_{GR} = \frac{1}{m^2} \int d^4x \sqrt{-g} \times \quad (1.7)$$

$$\times [\beta^2 R + (3 - 4\alpha)|\partial\beta|^2 + 16\pi m^2 L_M].$$

One then checks this against some Weyl-Dirac actions. Thus, neglecting terms $W^{ab}W_{ab}$ we find integrands involving $d^4x \sqrt{-g}$ times

$$-\beta^2 R + 3(3\sigma + 2)|\partial\beta|^2 + 2\Lambda\beta^4 + \mathfrak{L}_M \quad (1.8)$$

(see e.g. [11, 12, 17, 25]); the term $2\Lambda\beta^4$ of weight -4 is added gratuitously (recall $\Pi(\sqrt{-g}) = 4$). Consequently, omitting the Λ term, (1.8) corresponds to (1.7) times m^2 for $\mathfrak{L}_M \sim \sim 16\pi L_M$ and **(1D)** $9\sigma + 4\alpha + 3 = 0$. Hence one can identify conformal GR (without Λ) with a Bohmian-Weyl-Dirac theory where conformal mass \hat{m} corresponds to quantum mass \mathfrak{M} .

REMARK 1.1. The origin of a β^4 term in (1.8) from \hat{S}_{GR} in (1.2) with a term $2\sqrt{-\hat{g}}\hat{\Lambda}$ in the integrand would seem to involve writing **(1E)** $2\sqrt{-\hat{g}}\hat{\Lambda} = 2\sqrt{-\hat{g}}\hat{\phi}^2\Omega^4\hat{\Lambda} = = 2\sqrt{-g}\beta^4\hat{\Lambda}/m^4$ so that Λ in (1.8) corresponds to $\hat{\Lambda}$. Normally one expects $\Lambda\sqrt{-g} \rightarrow \sqrt{-\hat{g}}\hat{\phi}^2\Lambda$ (cf. [2]) or perhaps $\Lambda \rightarrow \hat{\phi}^2\Lambda = \Omega^{-4}\Lambda = \hat{\Lambda}$. In any case the role and nature of a cosmological constant seems to still be undecided. ■

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References

1. Adler R., Bazin M., and Schiffer M. Introduction to general relativity. McGraw-Hill, 1975.
2. Allemandi G., Capone M., Capozziello S., and Francaviglia M. arXiv: hep-th/0409198.
3. Audretsch J. *Phys. Rev. D*, 1981, v. 24, 1470–1477; 1983, v. 27, 2872–2884.
4. Audretsch J., Gähler F., and Straumann N. *Comm. Math. Phys.*, 1984, v. 95, 41–51.
5. Audretsch J. and Lämmerzahl C. *Class. Quant. Gravity*, 1988, v. 5, 1285–1295.
6. Blaschke D. and Dabrowski M. arXiv: hep-th/0407078.
7. Bonal R., Quiros I., and Cardenas R. arXiv: gr-qc/0010010.
8. Brans C. and Dicke R. *Phys. Rev.*, 1961, v. 124, 925–935.
9. Canuto V., Adams P., Hsieh S., and Tsiang E. *Phys. Rev. D*, 1977, v. 16, 1643–1663.
10. Carroll R. Fluctuations, information, gravity, and the quantum potential. Springer, 2006.
11. Carroll R. On the quantum potential. Arima Publ., 2007.
12. Carroll R. arXiv: gr-qc/0705.3921.
13. Castro C. *Found. Phys.*, 1922, v. 22, 569–615; *Found. Phys. Lett.*, 1991, v. 4, 81–99; *Jour. Math. Phys.*, 1990, v. 31, 2633–2636; On dark energy, Weyl’s geometry, different derivations of the vacuum energy density and the Pioneer anomaly. *Found. Phys.*, 2007, v. 37, no. 3, 366–409.
14. Castro C. and Mahecha J. *Prog. Phys.*, 2006, v. 1, 38–45.
15. Dabrowski M., Denkiewicz T., and Blaschke D. arXiv: hep-th/0507068.
16. Dirac P.A.M. *Proc. Royal Soc. London A*, 1951, v. 209, 291–296; 1952, v. 212, 330–339; 1973, v. 332, 403–418.
17. Israelit M. The Weyl-Dirac theory and our universe. Nova Science Publ., 1999.
18. Israelit M. *Found. Phys.*, 1998, v. 28, 205–228; 1999, v. 29, 1303–1322; 2002, v. 32, 295–321 and 945–961; arXiv: gr-qc/9608035.
19. Israelit M. and Rosen N. *Found. Phys.*, 1992, v. 22, 555–568; 1994, v. 24, 901–915; 1995, v. 25, 763.
20. Israelit M. *Gen. Relativity and Gravitation*, 1997, v. 29, 1411–1424 and 1597–1614.
21. Noldus J. arXiv: gr-qc/0508104.
22. Quiros I. *Phys. Rev. D*, 2000, v. 61, 124026; arXiv: hep-th/0009169.
23. Rosen N. *Found. Phys.*, 1982, v. 12, 213–224; 1983, v. 13, 363–372.
24. Santamato E. *Phys. Rev. D*, 1984, v. 29, 216–222; 1985, v. 32, 2615–2621.
25. Shojai F. and Shojai A. arXiv: gr-qc/0306099 and gr-qc/0404102.
26. Shojai F., Shojai A., and Golshani M. *Mod. Phys. Lett. A*, 1998, v. 13, 2725, 2915, and 2965.
27. Shojai F. and Golshani M. *Inter. Jour. Mod. Phys. A*, 1998, v. 13, 677–693 and 2135–2144.
28. Wheeler J. *Phys. Rev. D*, 1990, v. 41, 431–441; arXiv: hep-th/9706215, hep-th/9708088, hep-th/0002068, hep-th/0305017, gr-qc/9411030.